

Microscopically Calculated Shell-Model Effective TBMEs in the sd-Shell

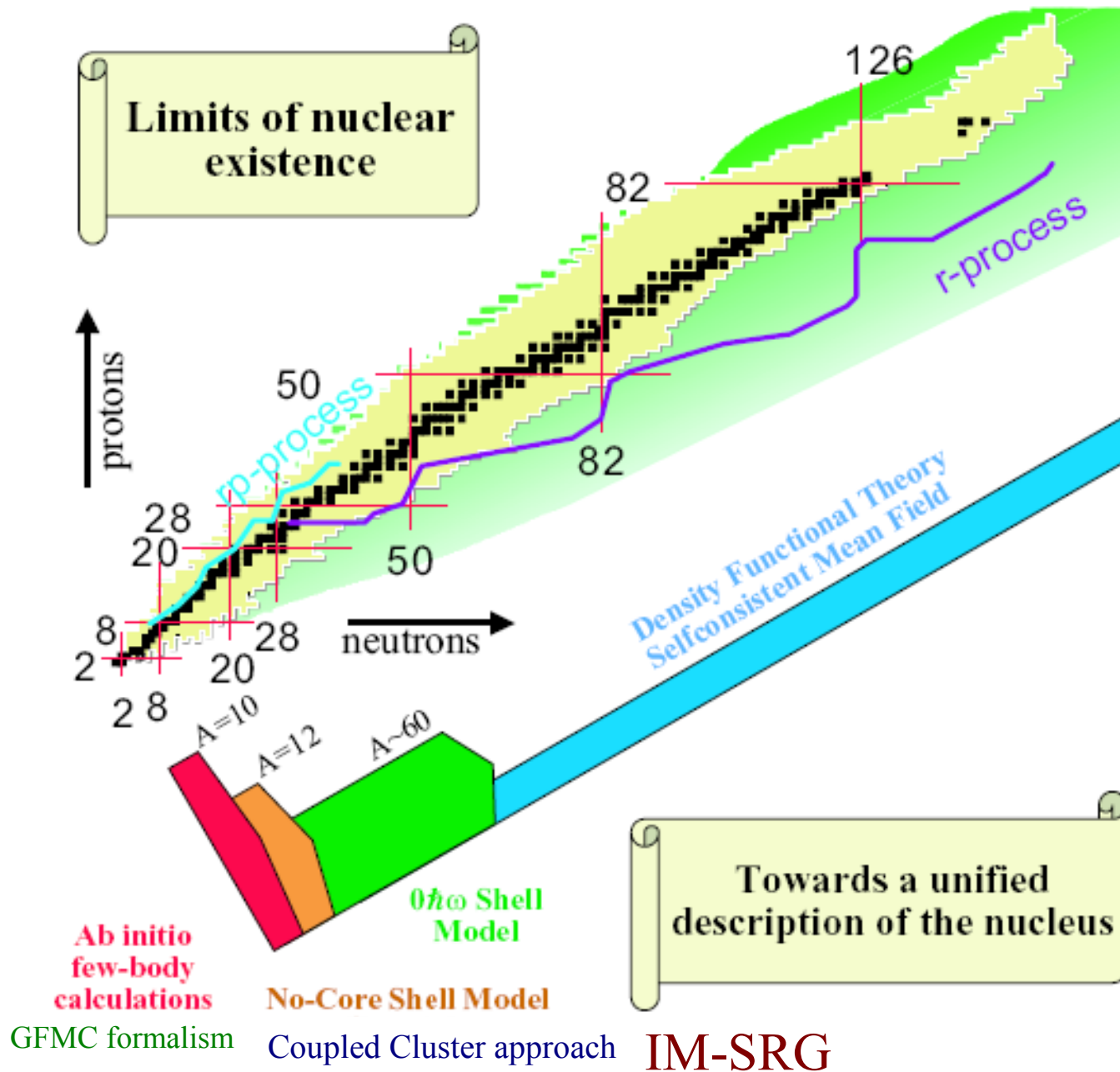
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OUTLINE

I. Overview of the *Ab Initio* Shell Model with a Core Approach

II. Results:

a.) General sd-shell

b.) Fluorine isotopes

c.) Effective Interactions Comparisons

III. Summary/Outlook

I. Overview of the *Ab Initio* Shell Model with a Core Approach

From few-body to many-body

Using the NCSM to calculate the shell model input

Ab initio
No Core Shell Model

Realistic NN & NNN forces

Effective interactions in
cluster approximation

Diagonalization of
many-body Hamiltonian

Core Shell Model

effective interactions for
valence nucleons

Diagonalization of the
Hamiltonian for valence
nucleons

Many-body experimental data



Effective interaction in a projected model space

$$H\Psi_\alpha = E_\alpha\Psi_\alpha \quad \text{where} \quad H = \sum_{i=1}^A t_i + \sum_{i<j}^A v_{ij}.$$

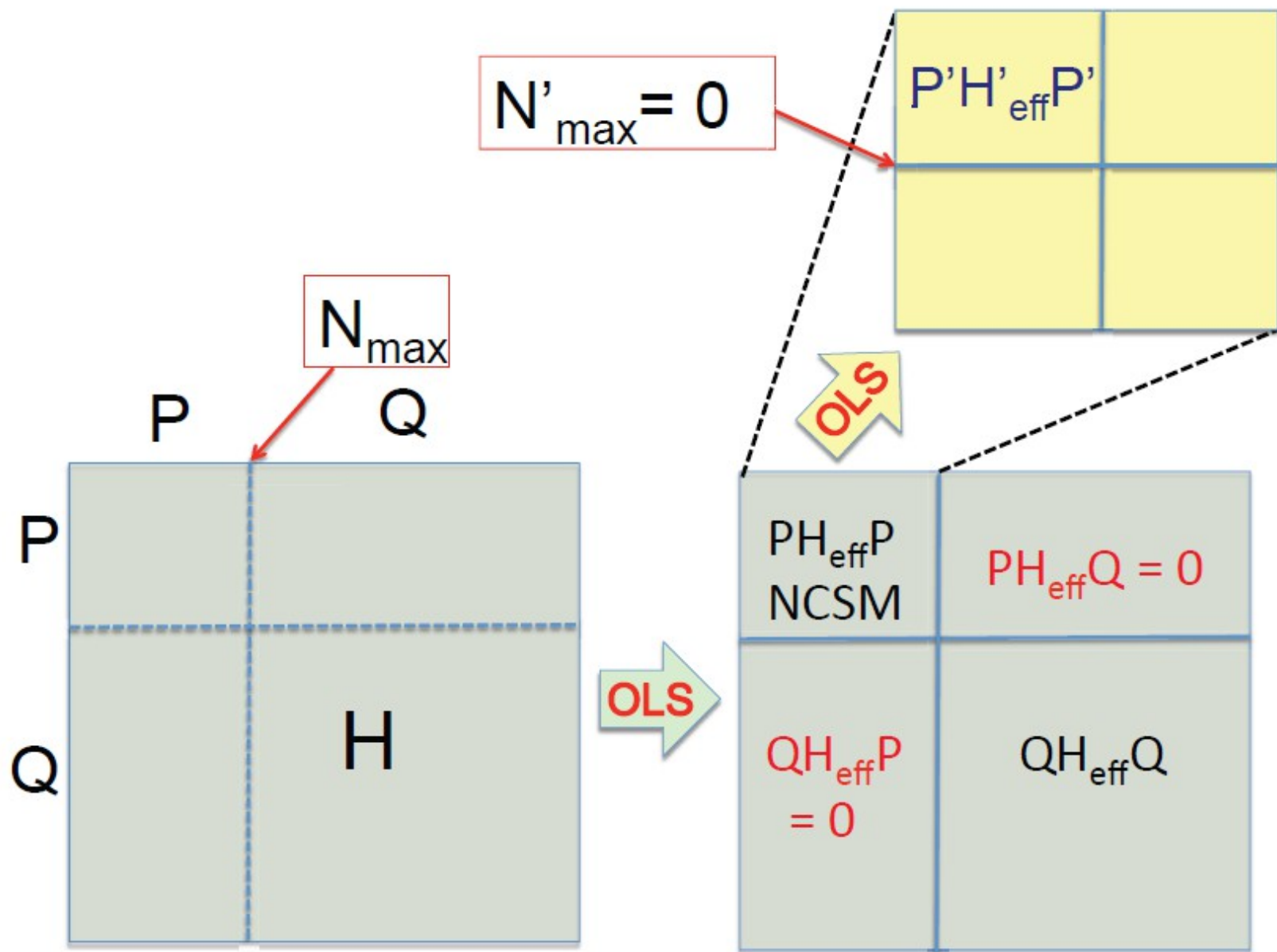
$$\mathcal{H}\Phi_\beta = E_\beta\Phi_\beta$$

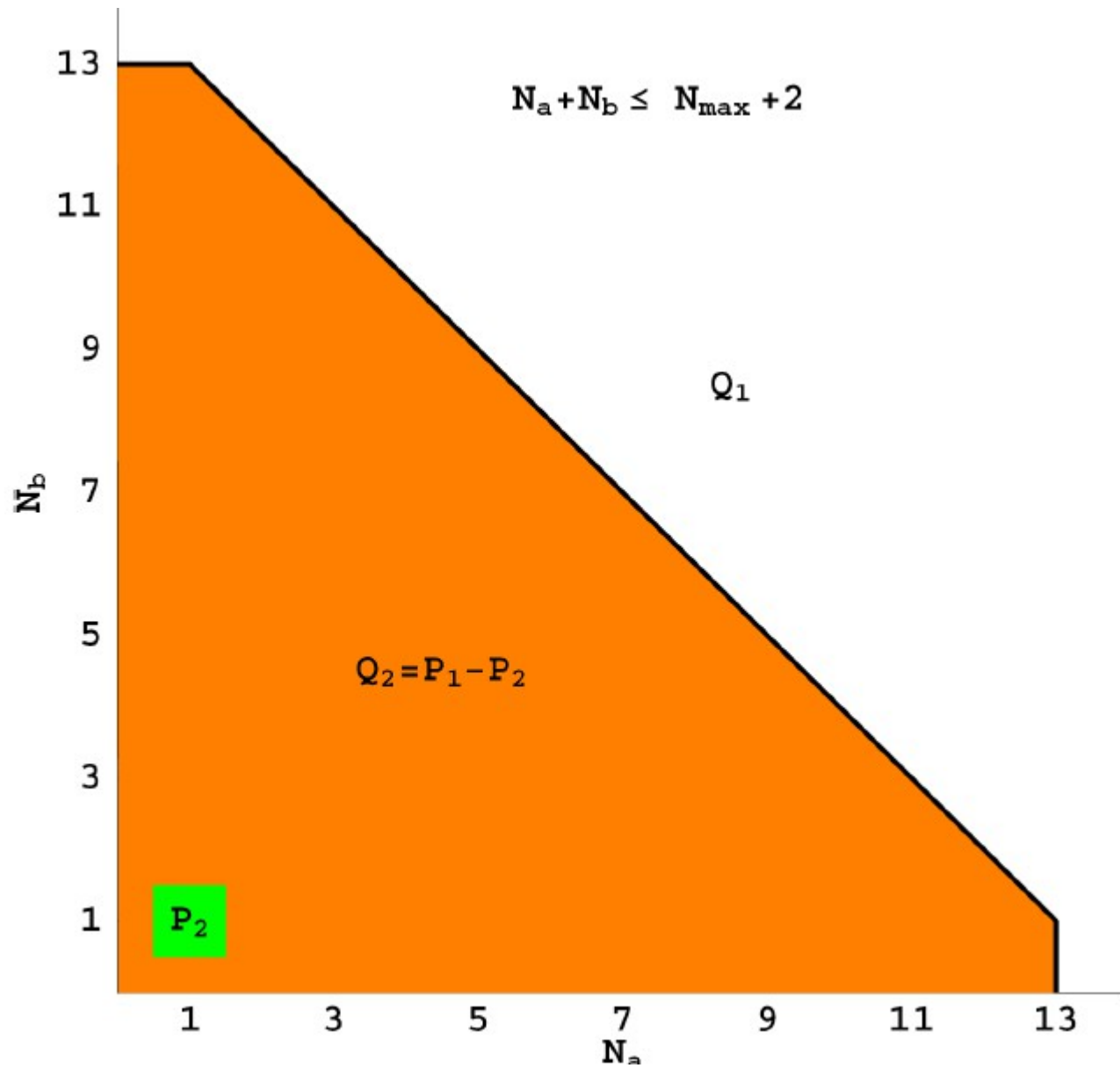
$$\Phi_\beta = P\Psi_\beta$$

P is a projection operator from S into \mathcal{S}

$$\langle \tilde{\Phi}_\gamma | \Phi_\beta \rangle = \delta_{\gamma\beta}$$

$$\mathcal{H} = \sum_{\beta \in \mathcal{S}} |\Phi_\beta\rangle E_\beta \langle \tilde{\Phi}_\beta|$$





PHYSICAL REVIEW C 78, 044302 (2008)

Ab-initio shell model with a core

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We construct effective two- and three-body Hamiltonians for the p -shell by performing $12\hbar\Omega$ *ab initio* no-core shell model (NCSM) calculations for $A = 6$ and 7 nuclei and explicitly projecting the many-body Hamiltonians onto the $0\hbar\Omega$ space. We then separate these effective Hamiltonians into inert core, one- and two-body contributions (also three-body for $A = 7$) and analyze the systematic behavior of these different parts as a function of the mass number A and size of the NCSM basis space. The role of effective three- and higher-body interactions for $A > 6$ is investigated and discussed.

DOI: [10.1103/PhysRevC.78.044302](https://doi.org/10.1103/PhysRevC.78.044302)

PACS number(s): 21.10.Hw, 21.60.Cs, 23.20.Lv, 27.20.+n

P. Navratil, M. Thoresen and B.R.B., Phys. Rev. C 55, R573 (1997)

FORMALISM

1. Perform a large basis NCSM for a core + 2N system, e.g., 18^{F} .
2. Use Okubo-Lee-Suzuki transformation to project these results into a single major shell to obtain effective 2-body matrix elements.
3. Separate these 2-body matrix elements into a core term, single-particle energies and residual 2-body interactions, i.e., the standard input for a normal Shell Model calculation.
4. Use these values for performing SM calculations in that shell.

II. Results: a.) sd-shell nuclei

Phys. Rev. C 91, 064301 (2015)

Ab initio effective interactions for *sd*-shell valence nucleons

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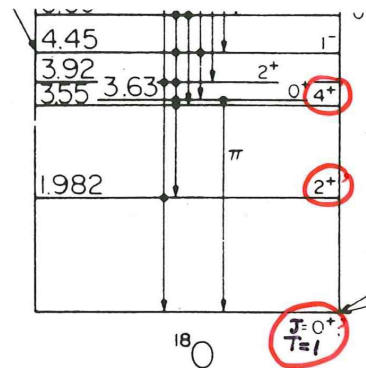
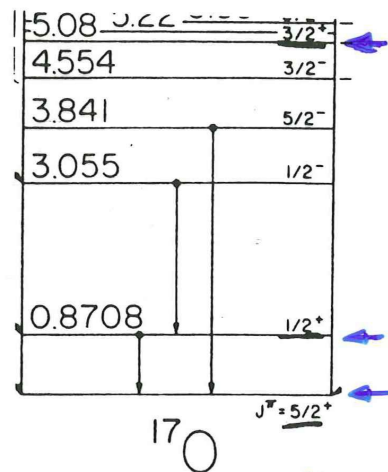
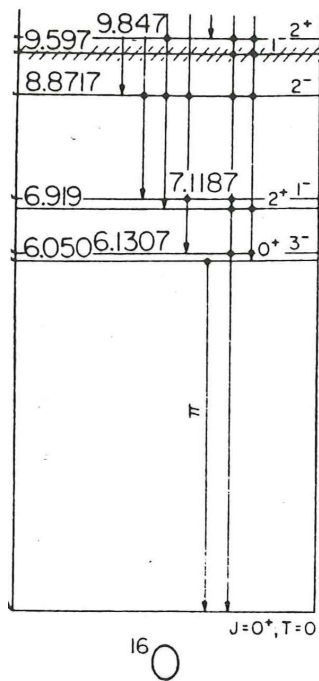
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(Dated: February 3, 2015)

We perform *ab initio* no core shell model calculations for $A = 18$ and 19 nuclei in a $4\hbar\Omega$, or $N_{\max} = 4$, model space using the effective JISP16 and chiral N3LO nucleon-nucleon potentials and transform the many-body effective Hamiltonians into the $0\hbar\Omega$ model space to construct the A -body effective Hamiltonians in the *sd*-shell. We separate the A -body effective Hamiltonians with $A = 18$ and $A = 19$ into inert core, one- and two-body components. Then, we use these core, one- and two-body components to perform standard shell model calculations for the $A = 18$ and $A = 19$ systems with valence nucleons restricted to the *sd*-shell. Finally, we compare the standard shell model results in the $0\hbar\Omega$ model space with the exact no core shell model results in the $4\hbar\Omega$ model space for the $A = 18$ and $A = 19$ systems and find good agreement.

ArXiv: Nucl-th 1502.00700

Empirical Single-Particle Energies



$$E_{0d_{5/2}} = 0.0 \text{ MeV}$$

$$E_{1s_{1/2}} = 0.87 \text{ MeV}$$

$$E_{0d_{3/2}} = 5.08 \text{ MeV}$$

$$H^{sd} (P \Phi)^{sd} = \left\{ \sum_i^{sd} \epsilon_i + V_{\text{eff}}^{sd} \right\} (P \Phi)^{sd}$$

$$\{H_0 + V_{\text{eff}}^{sd}\} (P \Phi)^{sd} = E^{sd} (P \Phi)^{sd}$$

Input: The results of $N_{\text{max}} = 4$ and $hw = 14$ MeV NCSM calculations

TABLE II: Proton and neutron single-particle energies (in MeV) for JISP16 effective interaction obtained for the mass of $A = 18$ and $A = 19$.

	$A = 18$			$A = 19$		
	$E_{\text{core}} = -115.529$			$E_{\text{core}} = -115.319$		
j_i	$\frac{1}{2}$	$\frac{5}{2}$	$\frac{3}{2}$	$\frac{1}{2}$	$\frac{5}{2}$	$\frac{3}{2}$
$\epsilon_{j_i}^n$	-3.068	-2.270	6.262	-3.044	-2.248	6.289
$\epsilon_{j_i}^p$	0.603	1.398	9.748	0.627	1.419	9.774

TABLE III: Proton and neutron single-particle energies (in MeV) for chiral N3LO effective interaction obtained for the mass of $A = 18$ and $A = 19$.

	$A = 18$			$A = 19$		
	$E_{\text{core}} = -118.469$			$E_{\text{core}} = -118.306$		
j_i	$\frac{1}{2}$	$\frac{5}{2}$	$\frac{3}{2}$	$\frac{1}{2}$	$\frac{5}{2}$	$\frac{3}{2}$
$\epsilon_{j_i}^n$	-3.638	-3.042	3.763	-3.625	-3.031	3.770
$\epsilon_{j_i}^p$	0.044	0.690	7.299	0.057	0.700	7.307

$A = 18$

Coupled Cluster, E_{core} : -130.462
Idaho NN N3LO + 3N N2LO

IM-SRG, E_{core} : -130.132
Idaho NN N3LO + 3N N2LO

$A = 19$

-130.056 from G.R. Jansen
et al. PRL 113,
142502 (2014)

-129.637 from H. Hergert
private comm.

No-Core Shell-Model Approach

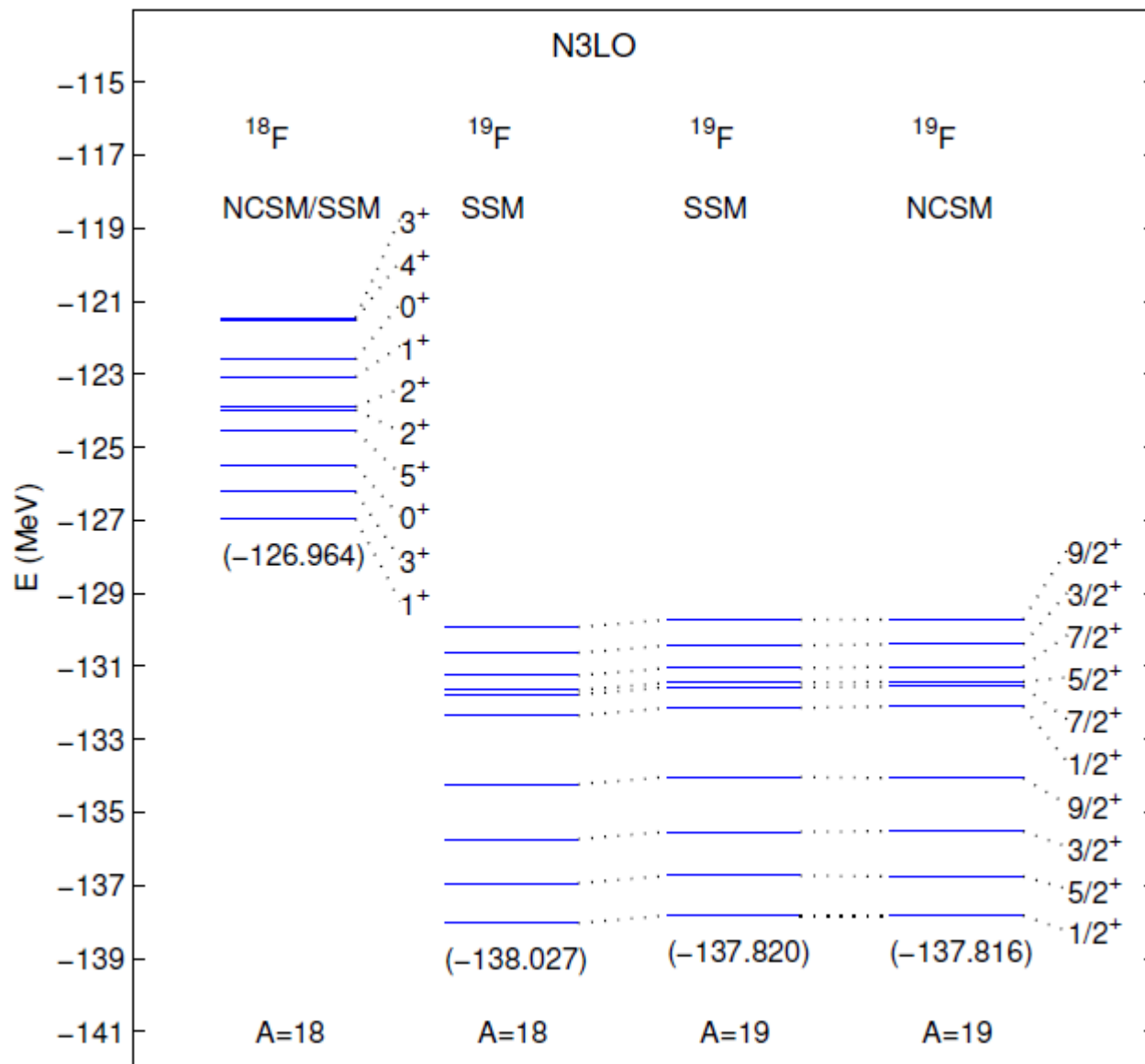
- Next, add CM harmonic-oscillator Hamiltonian

$$H_{CM}^{HO} = \frac{\vec{P}^2}{2Am} + \frac{1}{2}Am\Omega^2\vec{R}^2; \quad \vec{R} = \frac{1}{A}\sum_{i=1}^A \vec{r}_i, \quad \vec{P} = Am\dot{\vec{R}}$$

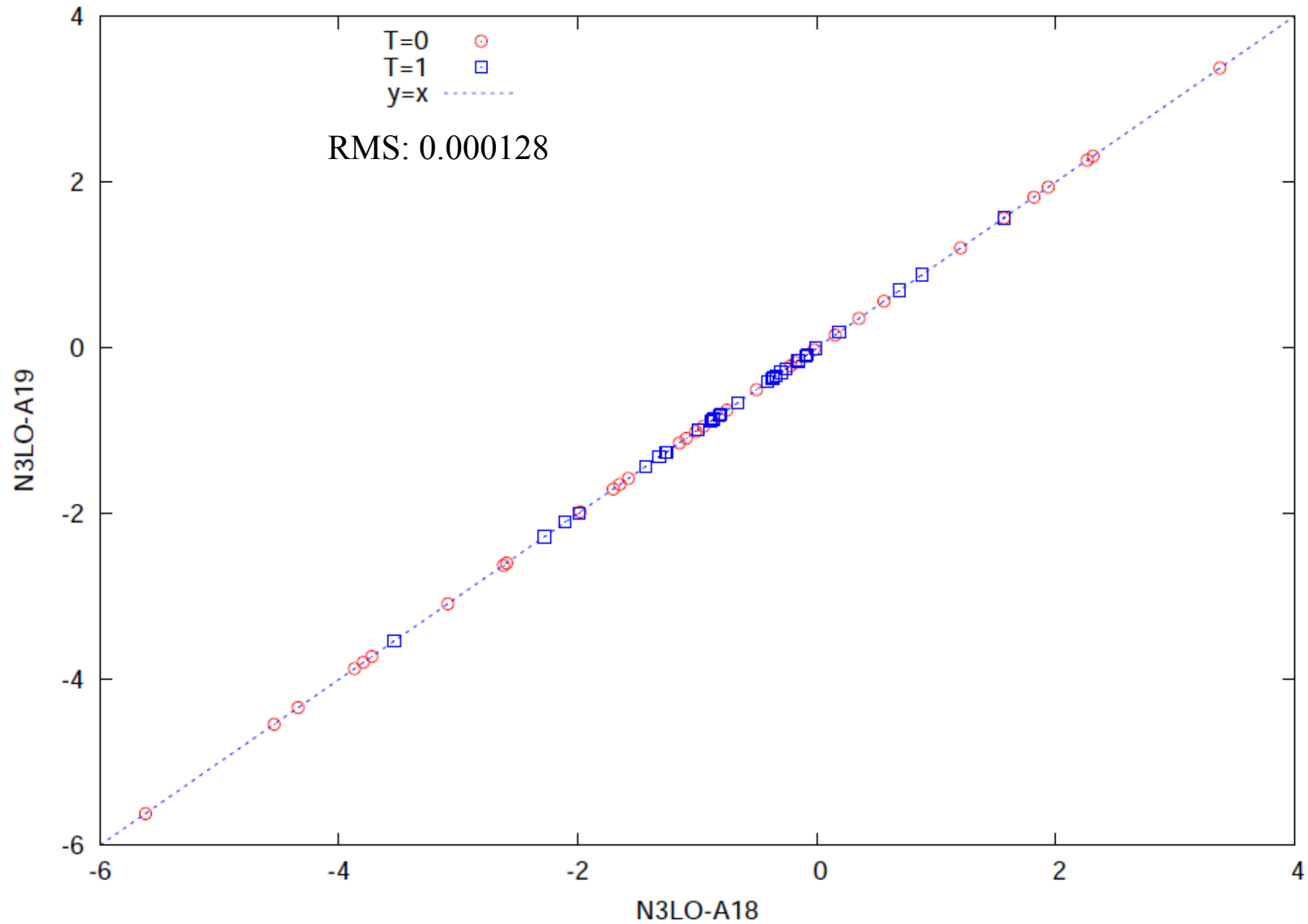
To H_A , yielding

$$H_A^\Omega = \sum_{i=1}^A \left[\frac{\vec{p}_i^2}{2m} + \frac{1}{2}m\Omega^2\vec{r}_i^2 \right] + \underbrace{\sum_{i<j=1}^A \left[V_{NN}(\vec{r}_i - \vec{r}_j) - \frac{m\Omega^2}{2A}(\vec{r}_i - \vec{r}_j)^2 \right]}_{V_{ij}}$$

Defines a basis (*i.e.* **HO**) for evaluating V_{ij}

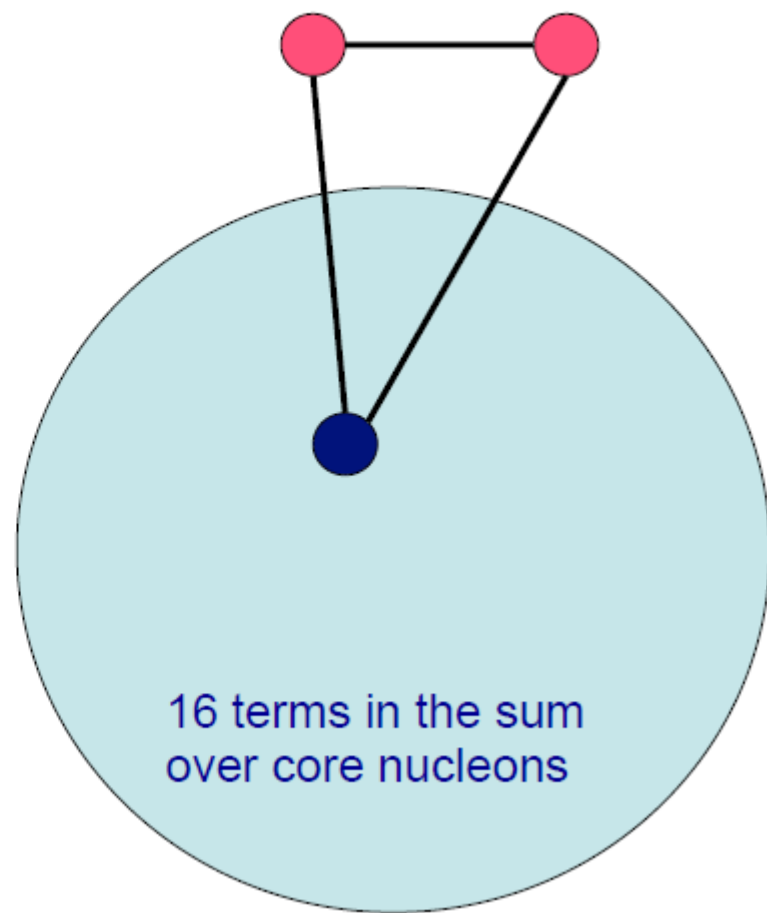
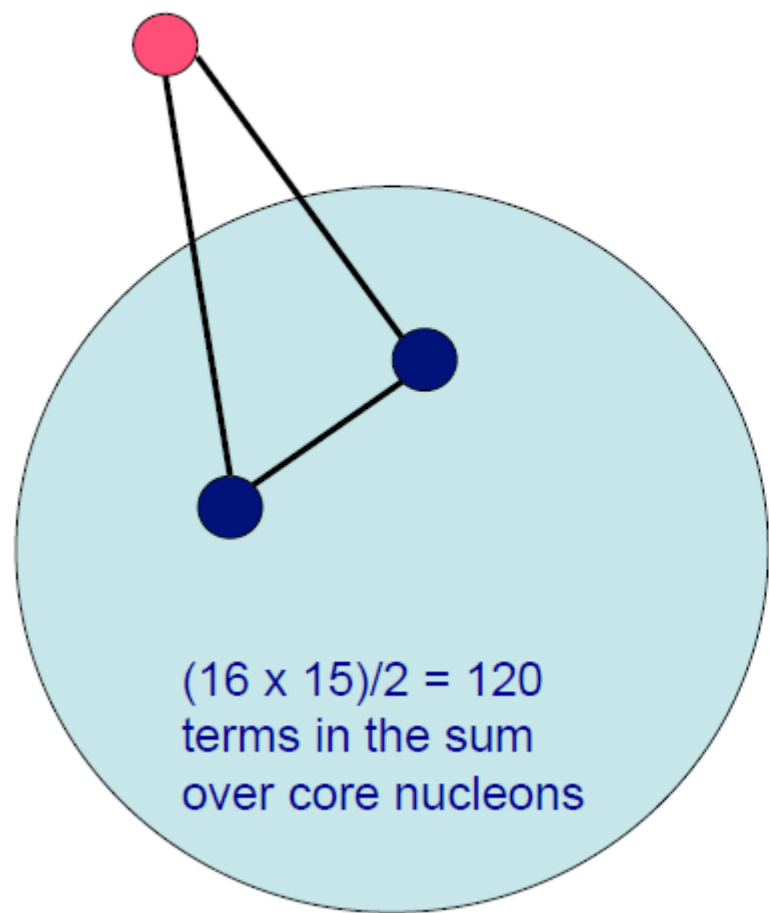


PRELIMINARY RESULT



II. Results: b) Fluorine isotopes

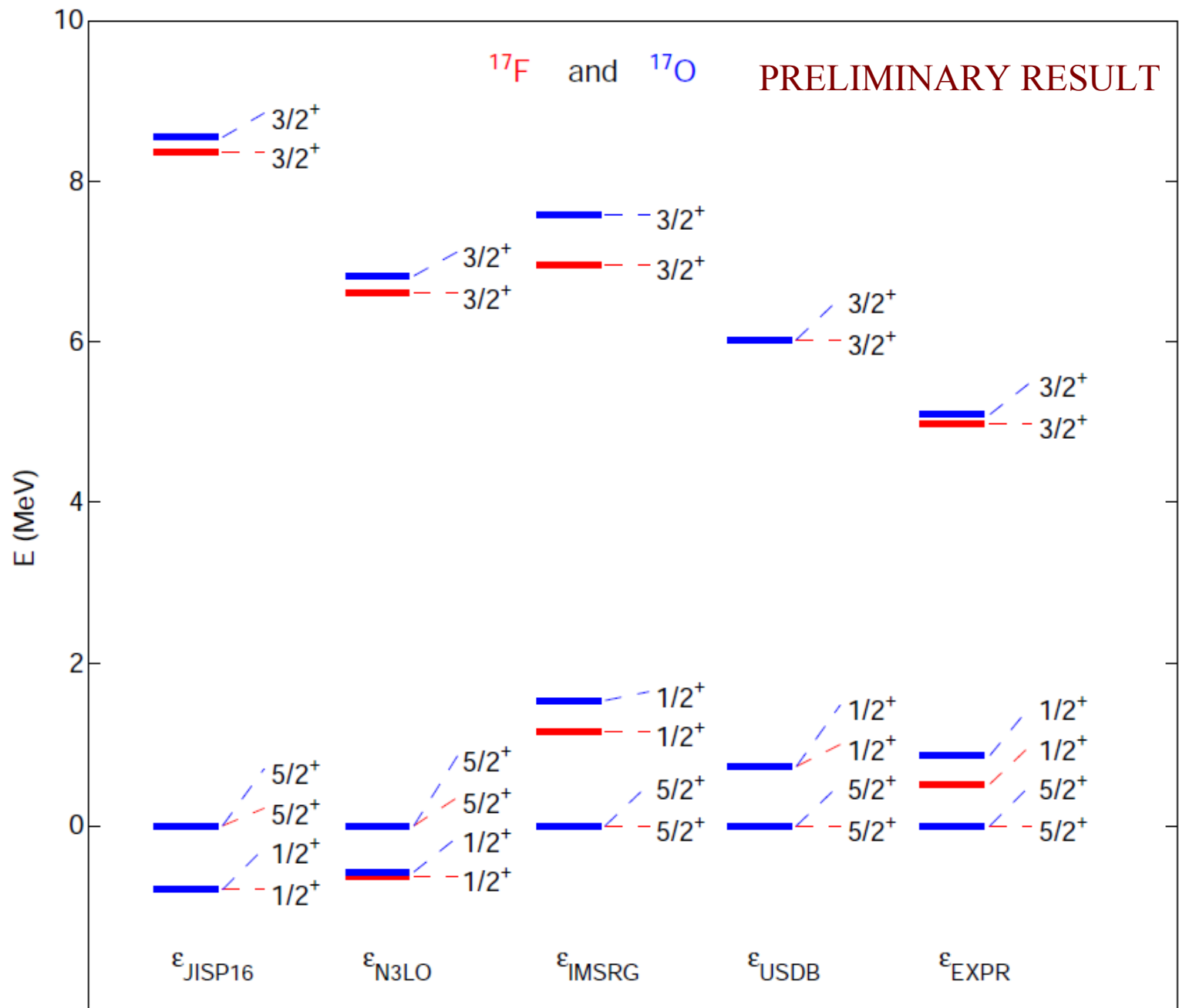
Effects of 3NFs on single-particle energies
vs effects of 3NFs on TBMEs



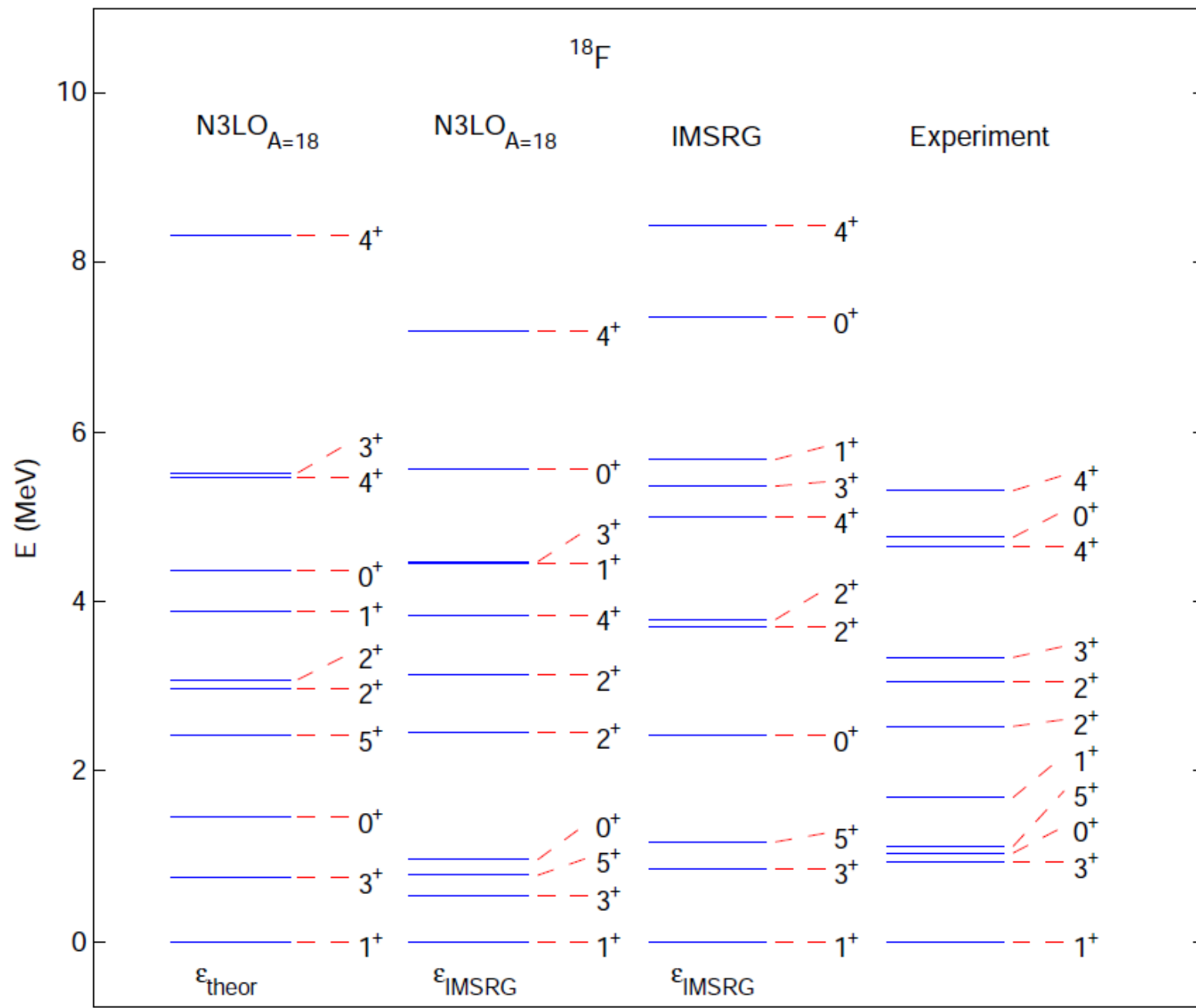
Survey of the Fluorine isotopes

1. Calculate the Fluorine isotopes using the same set of effective TBMEs, which are very weakly A -dependent, e.g., those determined from the N3LO NN interaction, to test how well they reproduce data trends.
2. Approximate the effect of 3NFs by replacing our theoretical single-particle energies with the theoretical ones obtained in the IM-SRG calculations of S.R. Stroberg et al.*
3. Compare our results for the F isotopes with those obtained with the IM-SRG approach* using an EFT N3LO NN plus N2LO NNN interaction and with experiment.

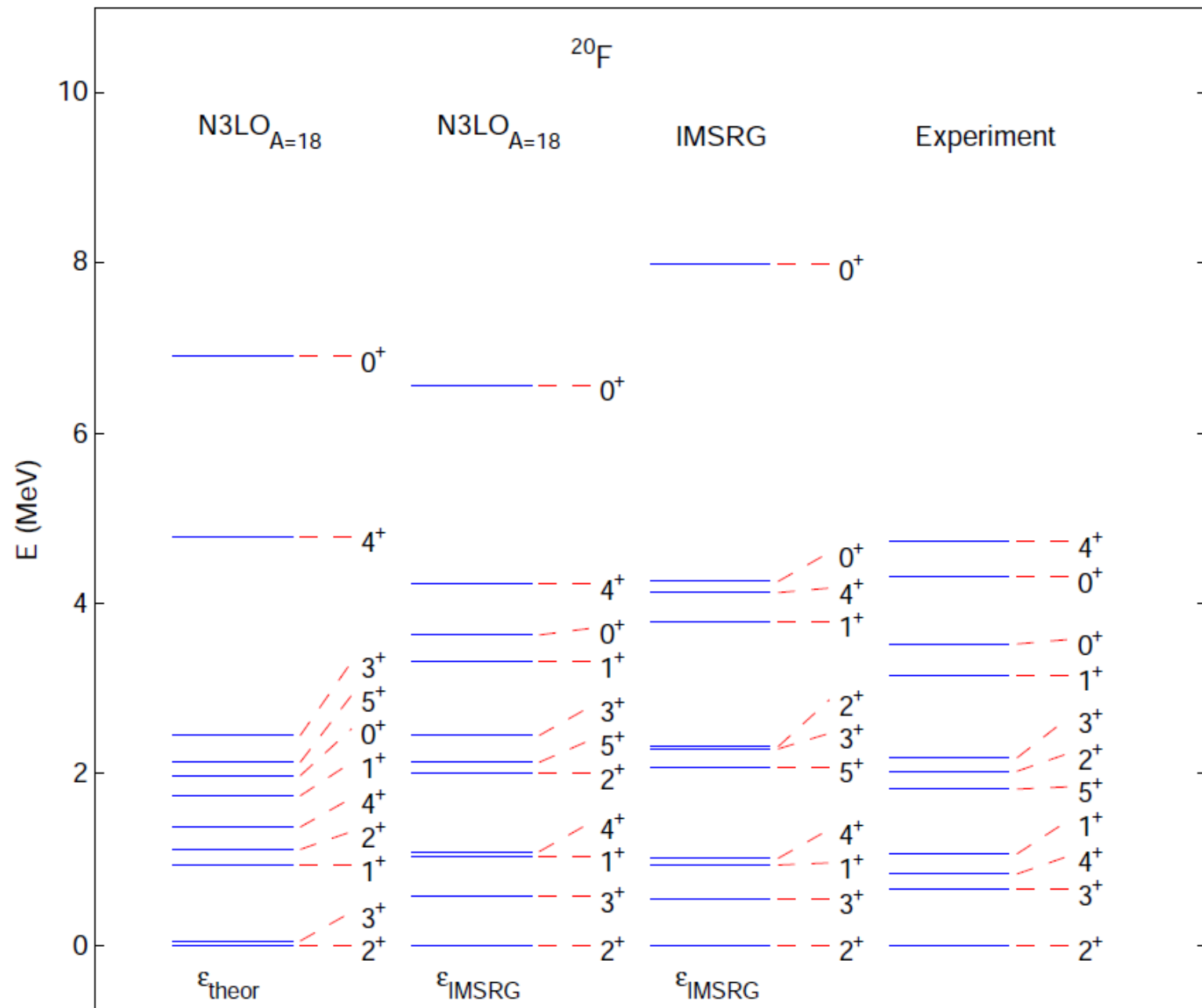
* S.R. Stroberg et al., arXiv Nucl-th 1511.02802 (2015)



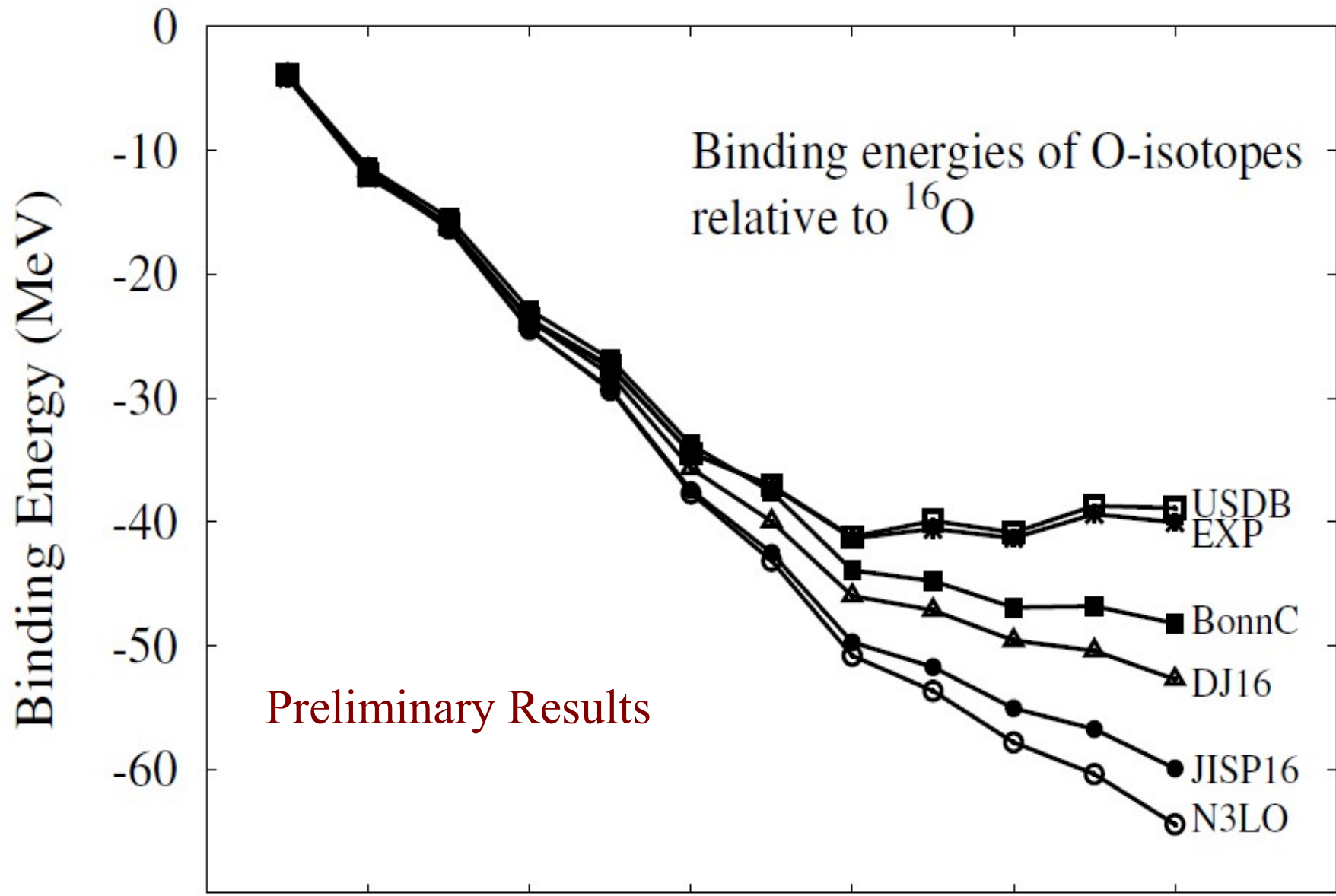
PRELIMINARY RESULT



PRELIMINARY RESULT

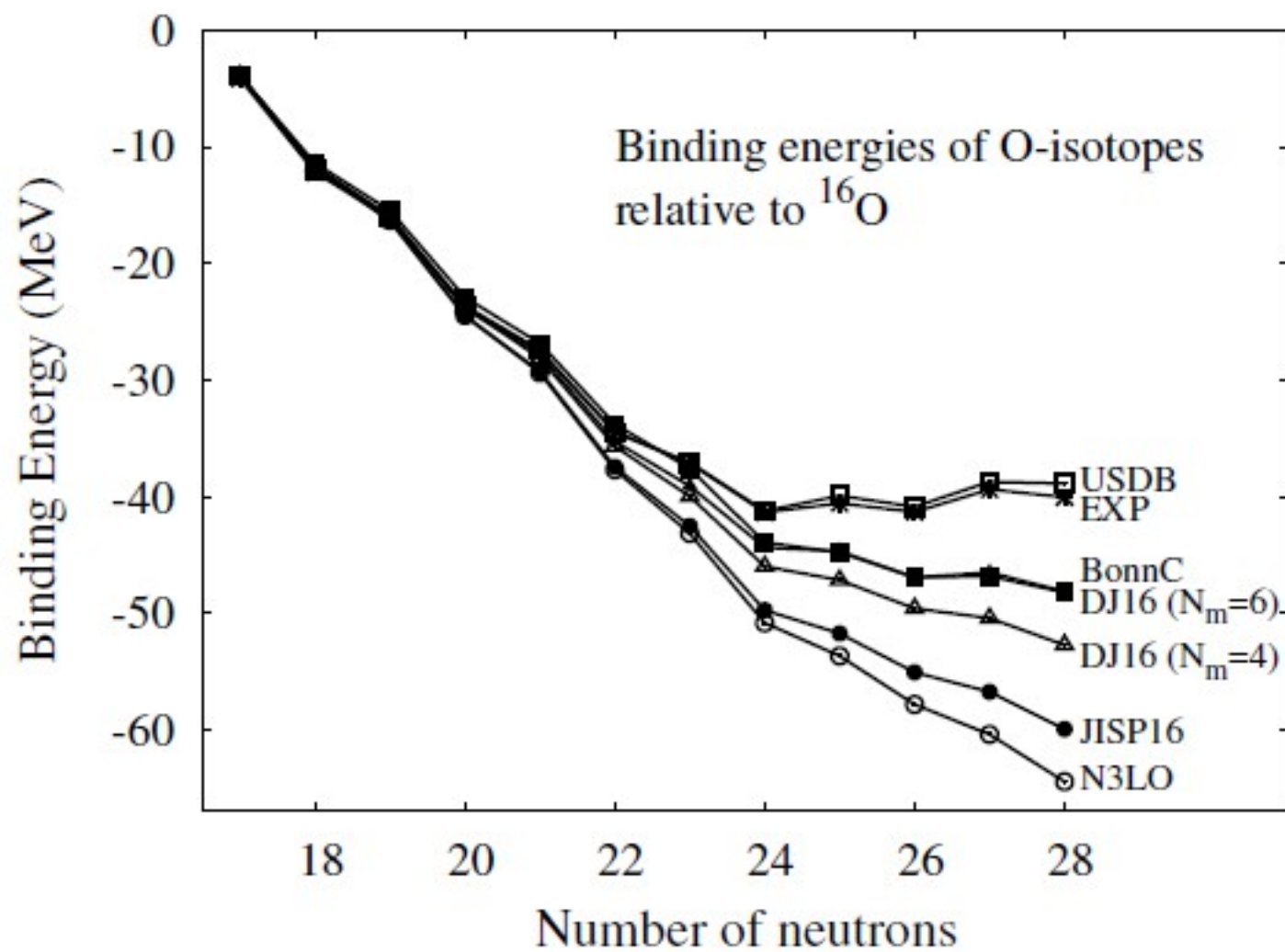


- II. Results: c) Effective Interaction Comparisons:
- i.) N3LO (E&M)
 - ii.) JISP16
 - iii.) DJ16



Points of Interest

1. The Monopole Term
2. The A-Dependence
3. Handcrafted Improvements



III. SUMMARY AND OUTLOOK

1. The interactions and approaches used in this study reproduced the gross trends and features of the experimental data for the Fluorine isotopes studied so far.
 2. Replacing our theoretical s.p. energies with those obtained in the IM-SRG calculations of Stroberg et al. to approximate the effects of a NNN interaction, in general, improved the agreement with experiment.
 3. The overall, reasonable agreement with experiment obtained using the IM-SRG approach with an EFT N3LO NN and N2LO NNN suggests that the trends in our results should continue to improve as we improve the interactions used and increase the size of our model space for our NCSM calculations.
 4. Of the three eff. Interactions: N3LO, JISP16 and DJ16; DJ16 is the most robust and closest to USDB in structure with adjustment.
- OUTLOOK:** Extend our calculations to other nuclei in the sd-shell to study charge dependence. Also study the hw dependence of our results.

Effective Hamiltonian for NCSM

Solving

$$\mathbf{H}_{A, a=2}^{\Omega} \Psi_{a=2} = \mathbf{E}_{A, a=2}^{\Omega} \Psi_{a=2}$$

in “infinite space” $2n+1 = 450$
relative coordinates

$P + Q = 1$; P – model space; Q – excluded space;

$$E_{A,2}^{\Omega} = U_2 H_{A,2}^{\Omega} U_2^{\dagger}$$

$$U_2 = \begin{pmatrix} U_{2,P} & U_{2,PQ} \\ U_{2,QP} & U_{2,Q} \end{pmatrix} \quad E_{A,2}^{\Omega} = \begin{pmatrix} E_{A,2,P}^{\Omega} & 0 \\ 0 & E_{A,2,Q}^{\Omega} \end{pmatrix}$$

$$H_{A,2}^{N_{\max}, \Omega, \text{eff}} = \frac{U_{2,P}^{\dagger}}{\sqrt{U_{2,P}^{\dagger} U_{2,P}}} E_{A,2,P}^{\Omega} \frac{U_{2,P}}{\sqrt{U_{2,P}^{\dagger} U_{2,P}}}$$

Two ways of convergence:

- 1) For $P \rightarrow 1$ and fixed a : $\widetilde{H}_{A, a=2}^{\text{eff}} \rightarrow H_A$
- 2) For $a \rightarrow A$ and fixed P : $\widetilde{H}_{A, a}^{\text{eff}} \rightarrow H_A$

Effective Hamiltonian for SSM

How to calculate the Shell Model 2-body effective interaction:

Two ways of convergence:

1) For $P \rightarrow 1$ and fixed a : $H_{A,a=2}^{\text{eff}} \rightarrow H_A$: previous slide

2) For $a_1 \rightarrow A$ and fixed P_1 : $H_{A,a_1}^{\text{eff}} \rightarrow H_A$

$P_1 + Q_1 = P$; P_1 - small model space; Q_1 - excluded space;

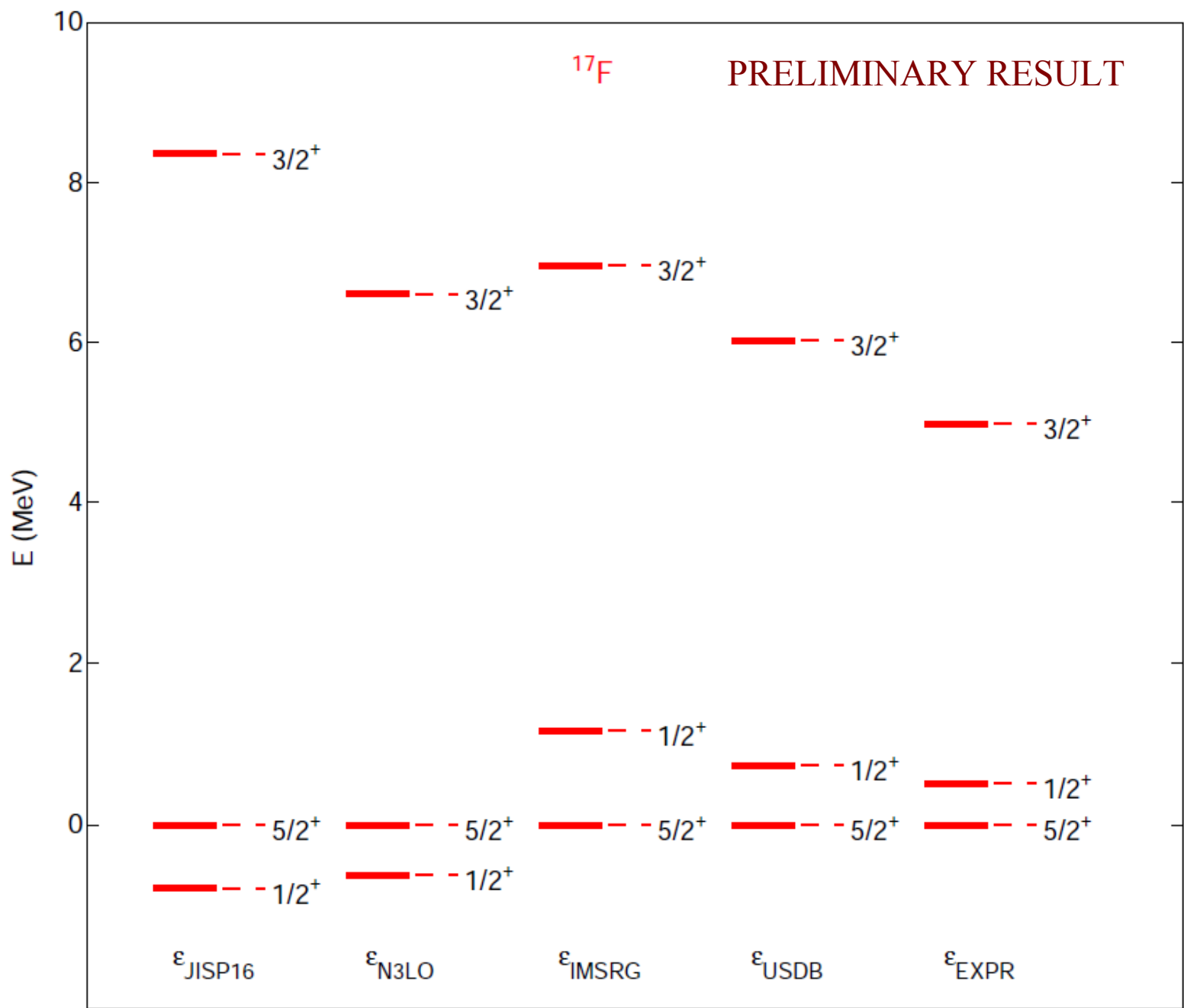
$$\mathcal{H}_{A,a_1}^{N_{1,\max}, N_{\max}} = \frac{U_{a_1, P_1}^{A, \dagger}}{\sqrt{U_{a_1, P_1}^{A, \dagger} U_{a_1, P_1}^A}} E_{A, a_1, P_1}^{N_{\max}, \Omega} \frac{U_{a_1, P_1}^A}{\sqrt{U_{a_1, P_1}^{A, \dagger} U_{a_1, P_1}^A}}$$

Valence Cluster Expansion

$N_{1,\max} = 0$ space (p-space); $a_1 = A_c + a_v$; a_1 - order of cluster;

A_c - number of nucleons in core; a_v - order of valence cluster;

$$\mathcal{H}_{A,a_1}^{0, N_{\max}} = \sum_k^{a_v} V_k^{A, A_c + k}$$



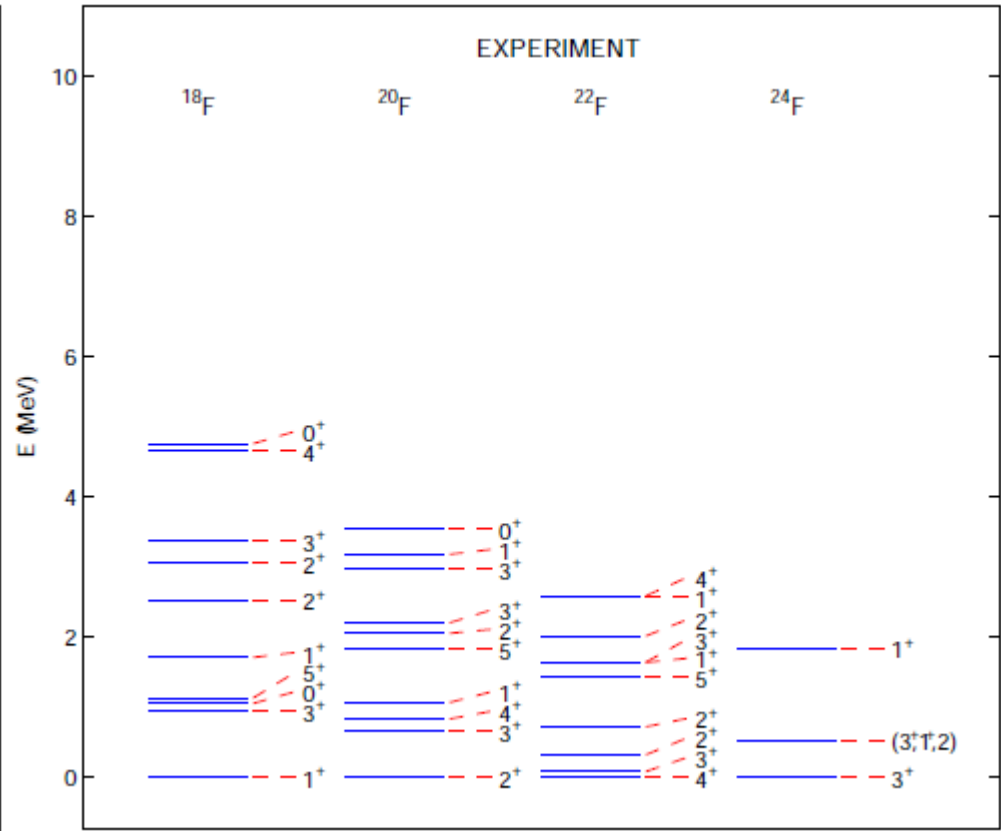
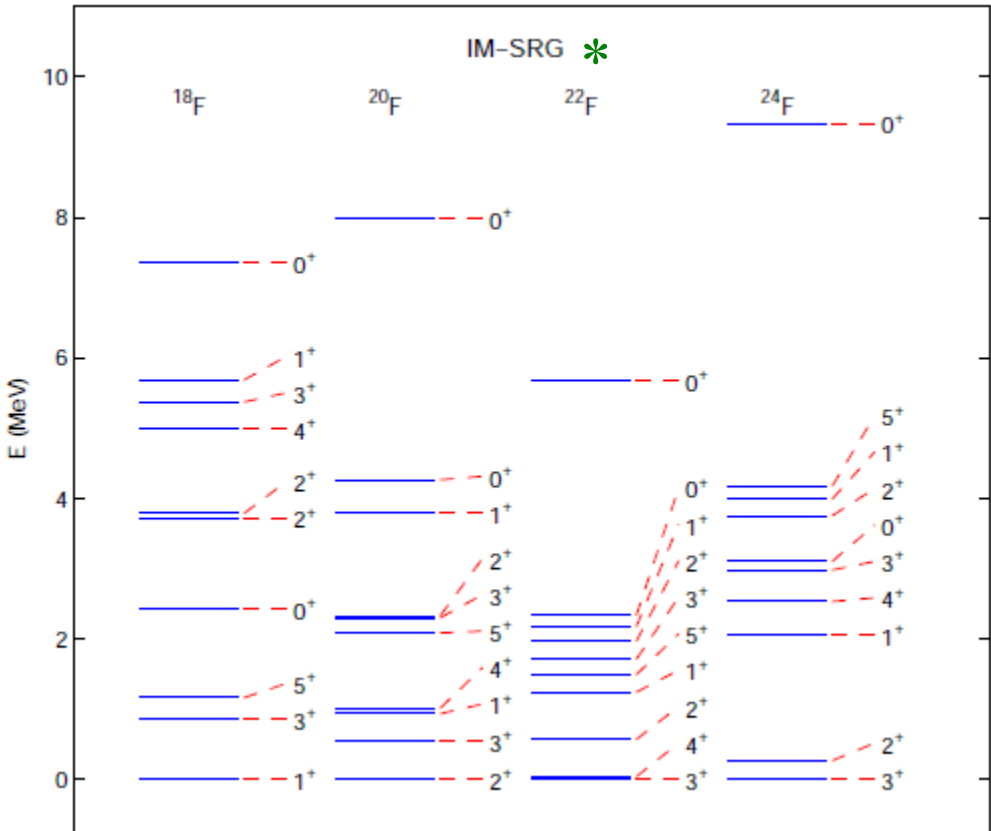
Preliminary Results

TABLE I: The single-particle energies (in MeV) used in the standard shell model calculations of F isotopes. (n) and (p) represent neutron and proton, respectively.

j_i	$\frac{1}{2}$	$\frac{5}{2}$	$\frac{3}{2}$	
JISP16 $_{A=18}$ (^{17}O) : (n)	-3.068	-2.270	6.262	
JISP16 $_{A=18}$ (^{17}F) : (p)	0.603	1.398	9.748	
USDA	-3.0612	-3.9436	1.9798	*
USDB	-3.2079	-3.9257	2.1117	
IM-SRG (^{17}O) : (n)	-3.089	-4.643	2.940	**
IM-SRG (^{17}F) : (p)	0.255	-0.909	6.035	

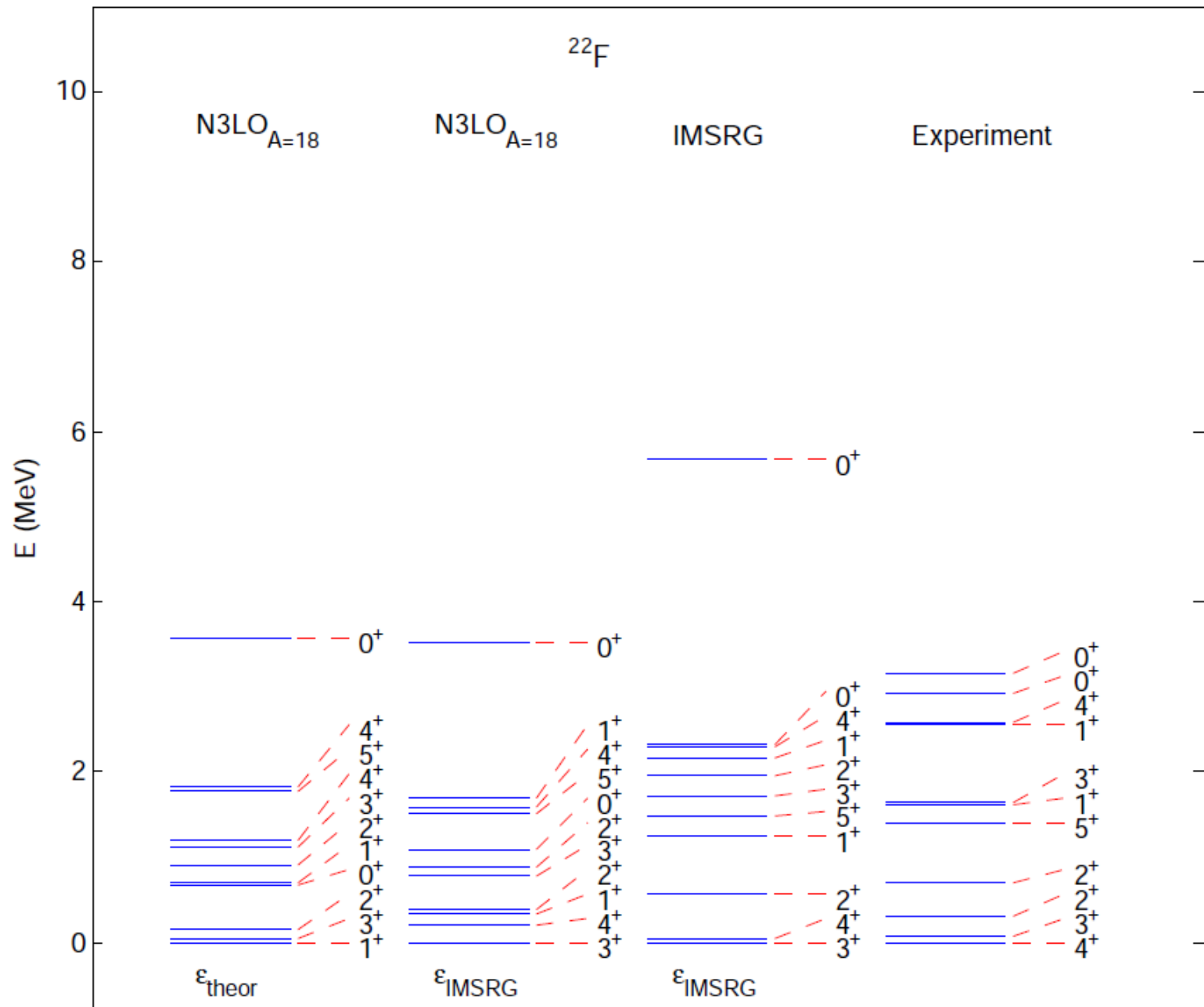
* B.A. Brown & W.A. Richter, PRC 74, 034315

** (2006) S.R. Stroberg, et al., arXiv Nucl-th 1511.02802

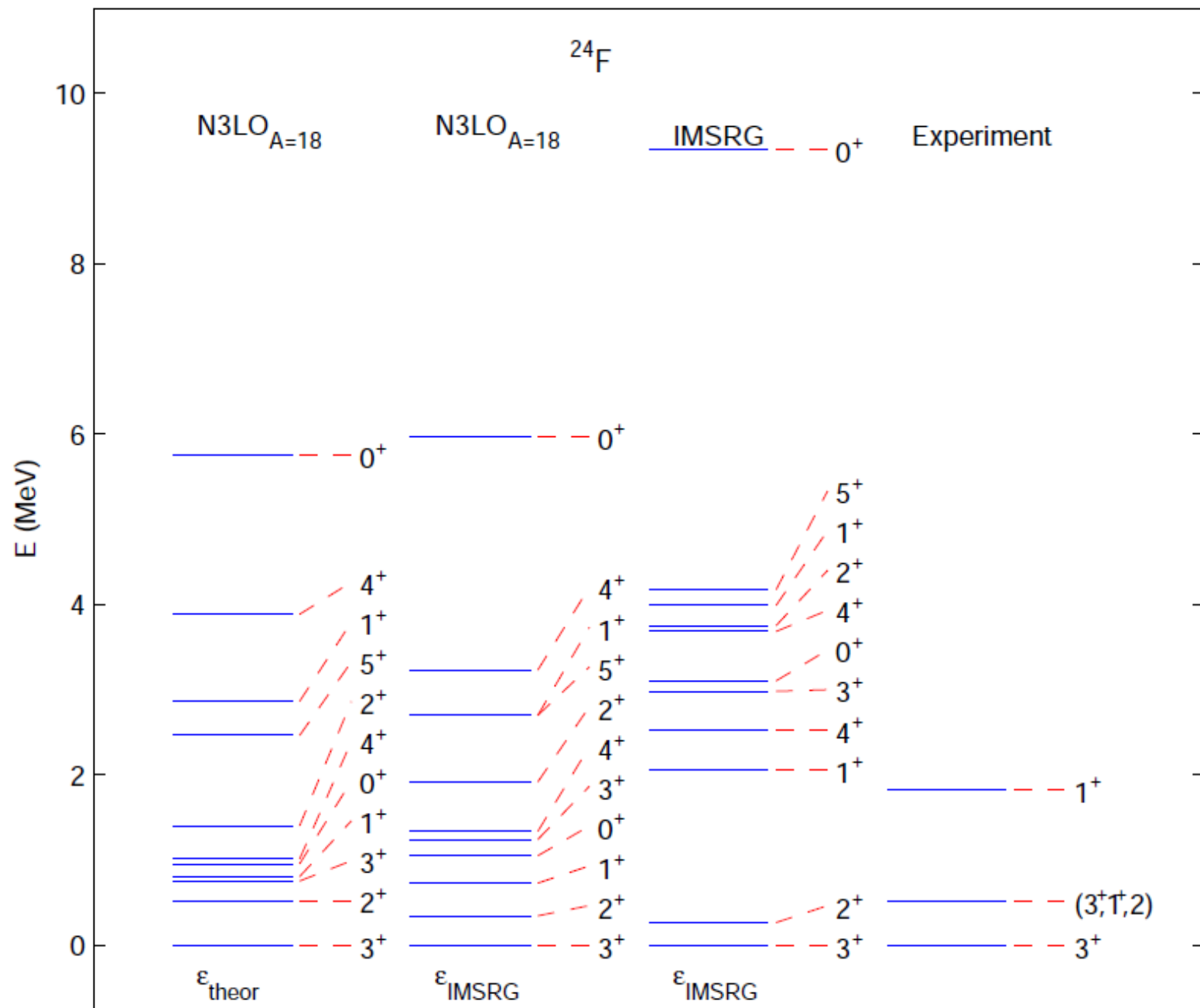


* S.R. Stroberg, et al., arXiv Nucl-th 1511.02802

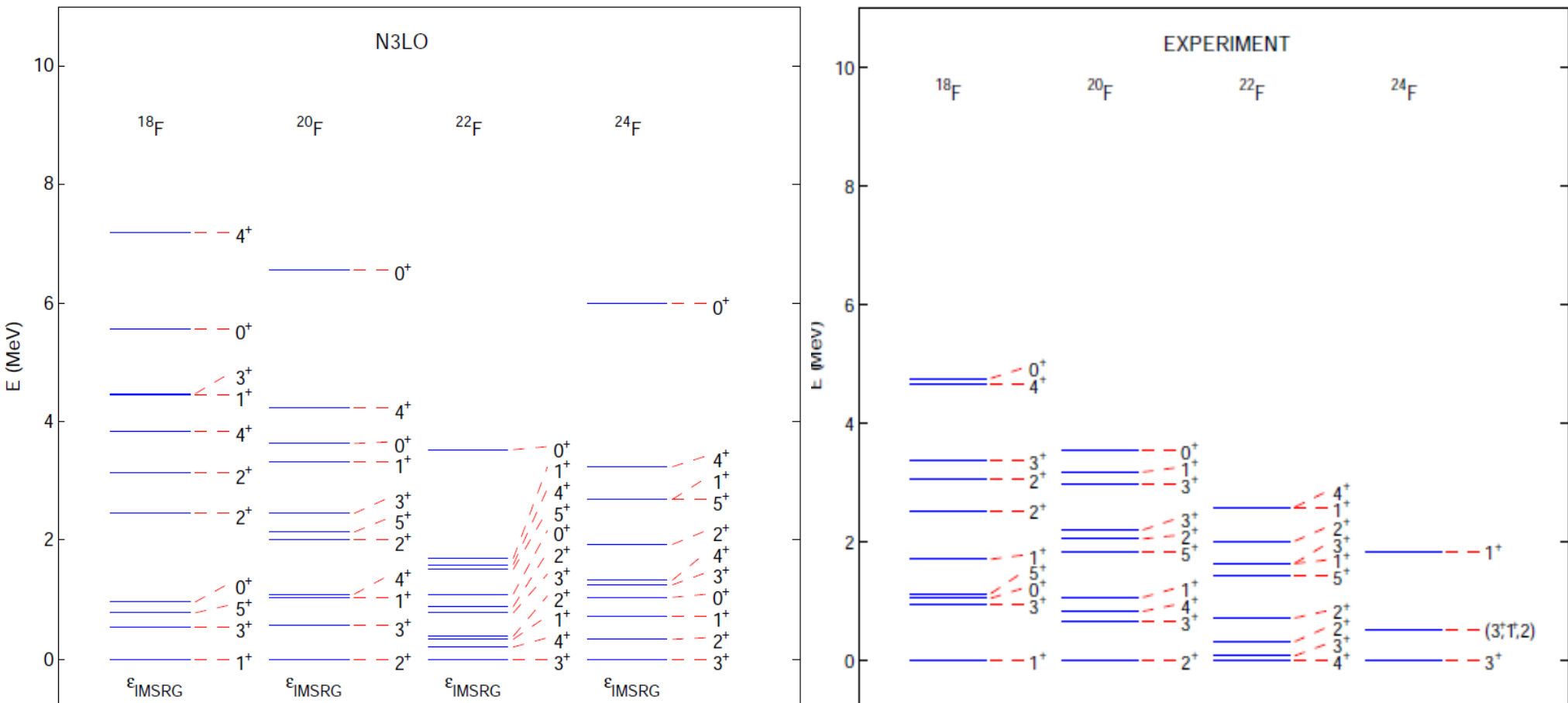
PRELIMINARY RESULT



PRELIMINARY RESULT

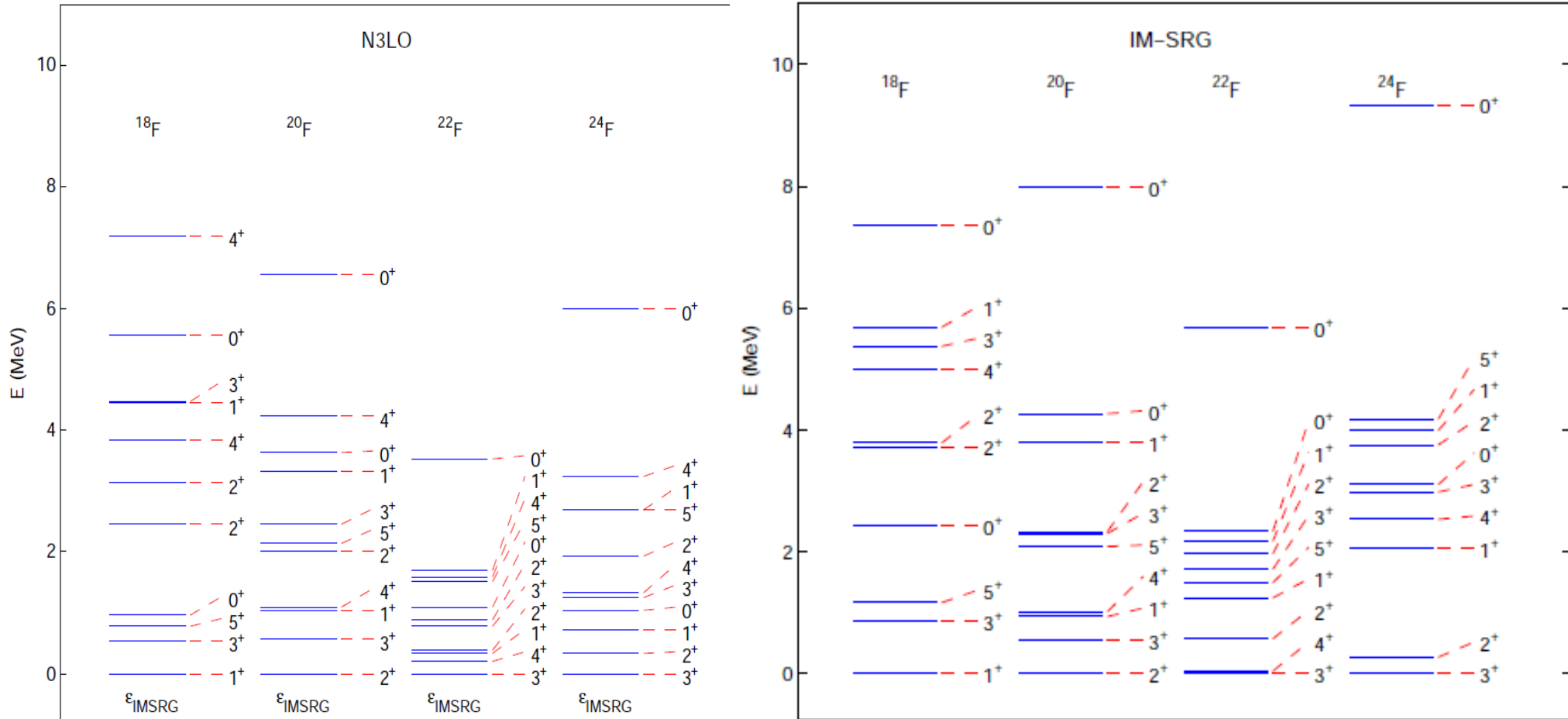


PRELIMINARY RESULTS



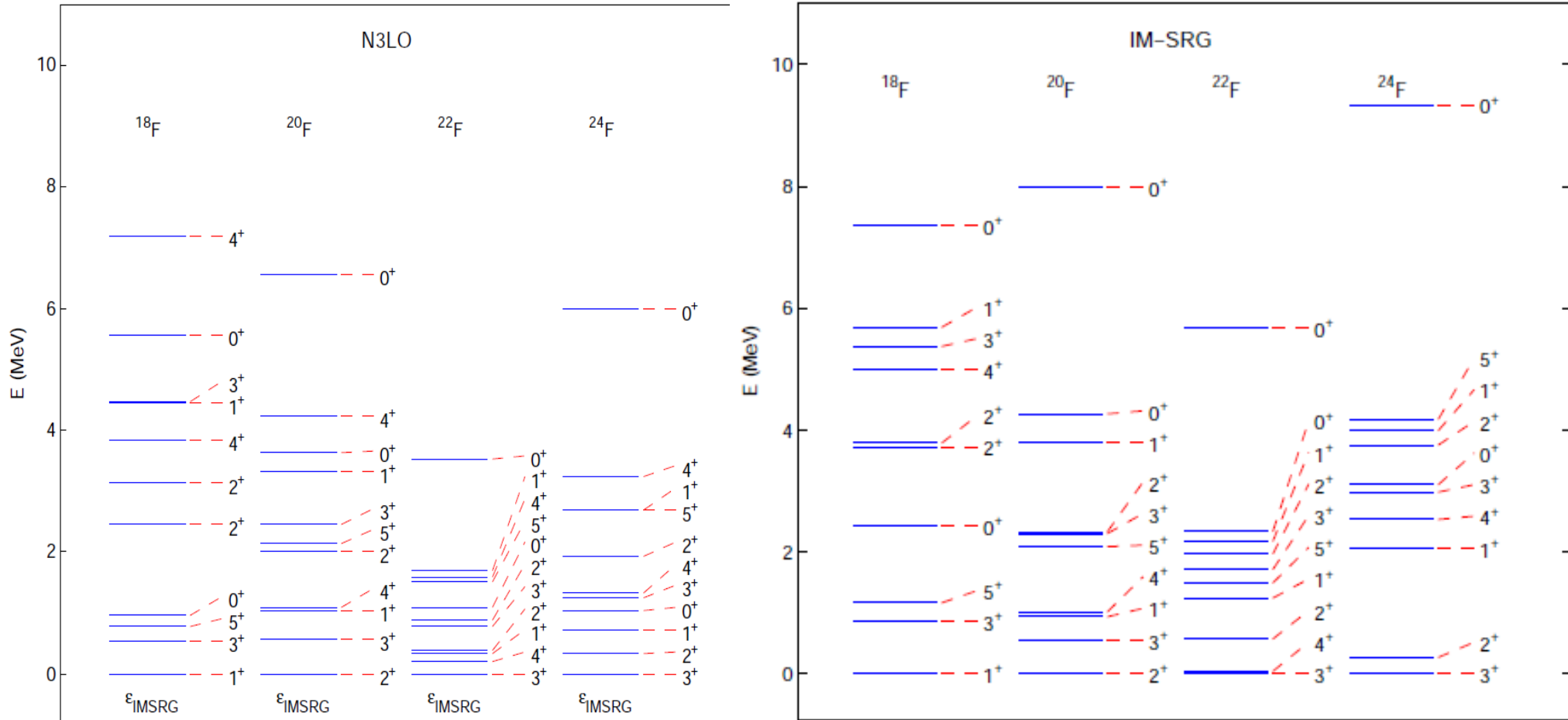
N3LO + IMSRG spe vs Experiment

PRELIMINARY RESULTS



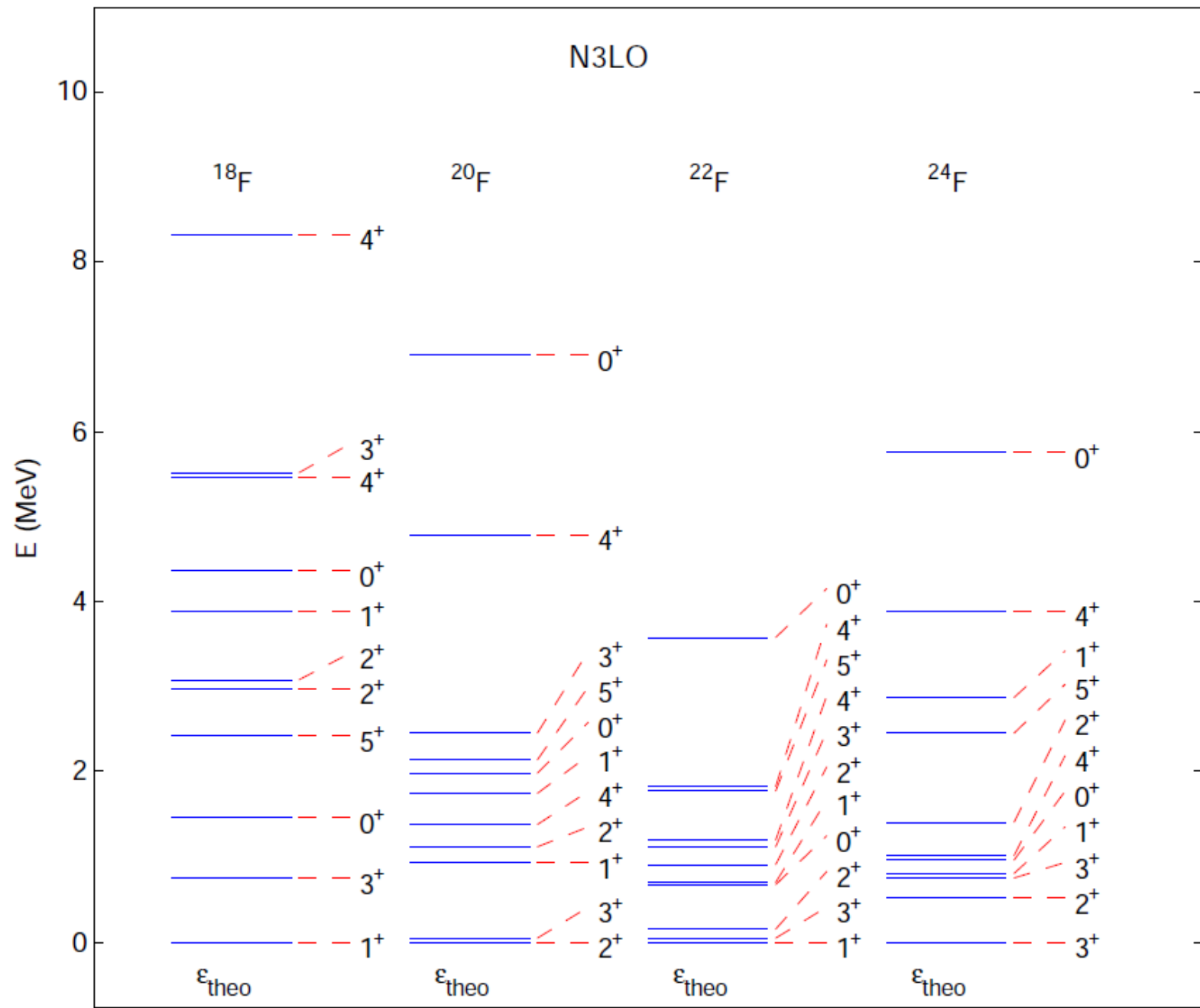
N3LO + IMSRG spe vs IM-SRG

PRELIMINARY RESULTS

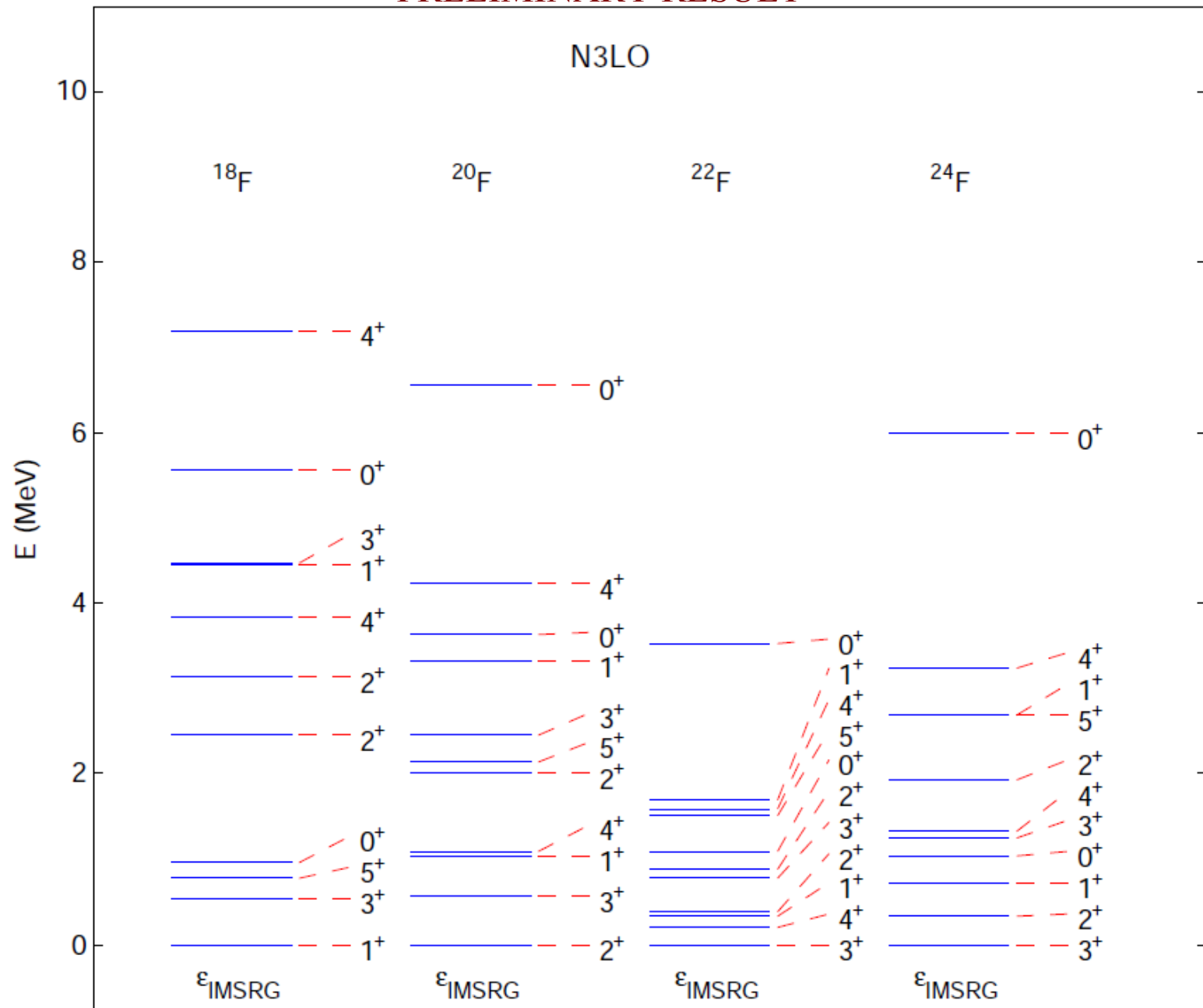


N3LO + IMSRG spe vs IM-SRG

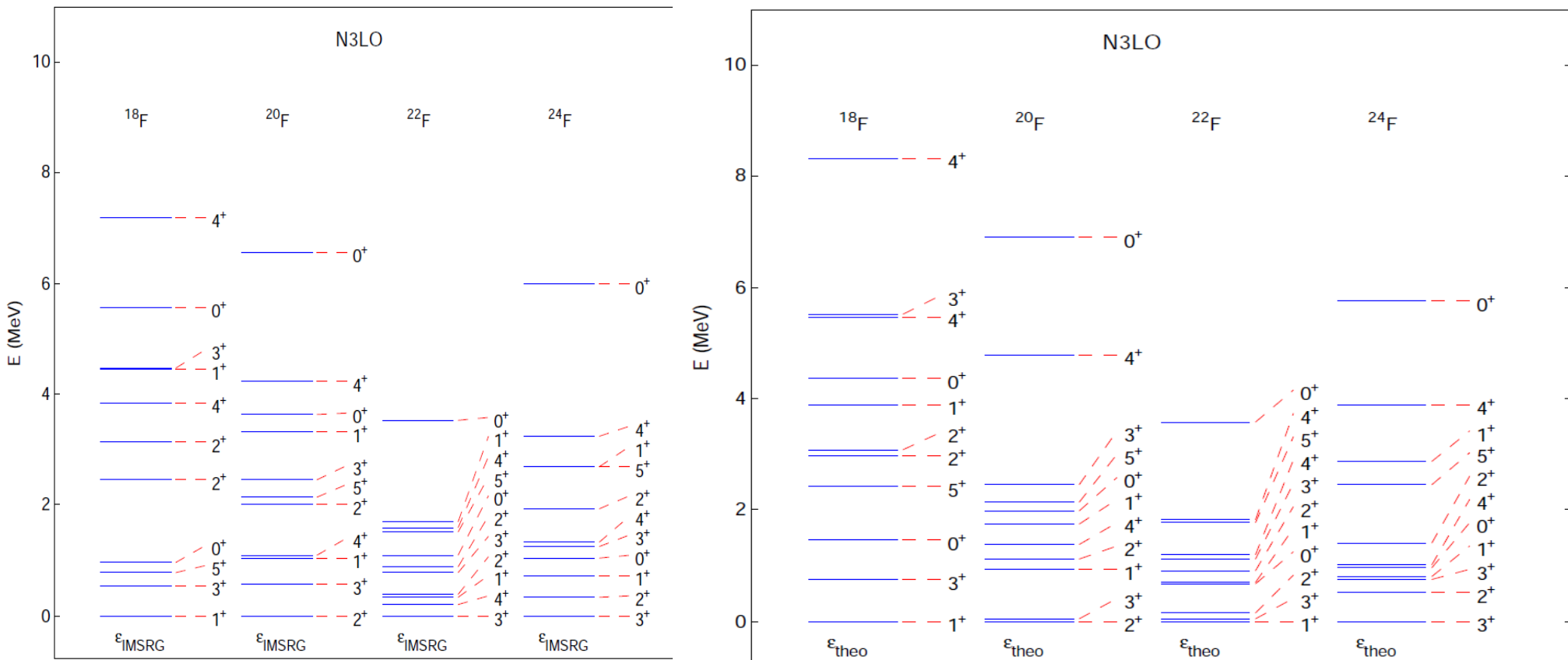
PRELIMINARY RESULT



PRELIMINARY RESULT



PRELIMINARY RESULTS



N3LO + IMSRG spe vs N3LO + theoretical spe